

Reactors and Fuels

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Major Topics

- Neutron interactions with matter and reactor control
- Nuclear reactors
- Nuclear fuels: fresh and spent

Interaction of Neutrons with Matter and Reactor Control

Interactions of Neutrons with Matter

Microscopic Neutron Cross Section

- Probability that a neutron will interact with a nucleus
- Measured in cm^2 , denoted by σ
- Cross sections are on the order of 10^{-24} cm^2 with a typical range of $\pm 100x$
- To early researchers this was unexpectedly large - as big as a barn – and so it is called
- The cross section is not a measurement of the physical cross section of the nucleus
 - More a measure of how much more stable the nucleus will be after the interaction occurs

Reaction Rate

- Macroscopic cross section: Microscopic cross section (cm^2) multiplied by density of target nuclide (N , atoms/cm^3)
 - Denoted by Σ
- When multiplied by the neutron flux (ϕ , $\text{neutrons}/\text{cm}^2\text{-sec}$) the result is the interaction rate (R , $\text{interactions}/\text{cm}^3\text{-sec}$)
- $R = N \cdot \sigma \cdot \phi = \Sigma \cdot \phi$

Neutron Flux

- The neutron flux is the total path length covered by all neutrons going in all directions in a cubic centimeter

$$\phi = n \cdot v$$

where

n = neutron density, $n/\text{cm}^3\text{-sec}$

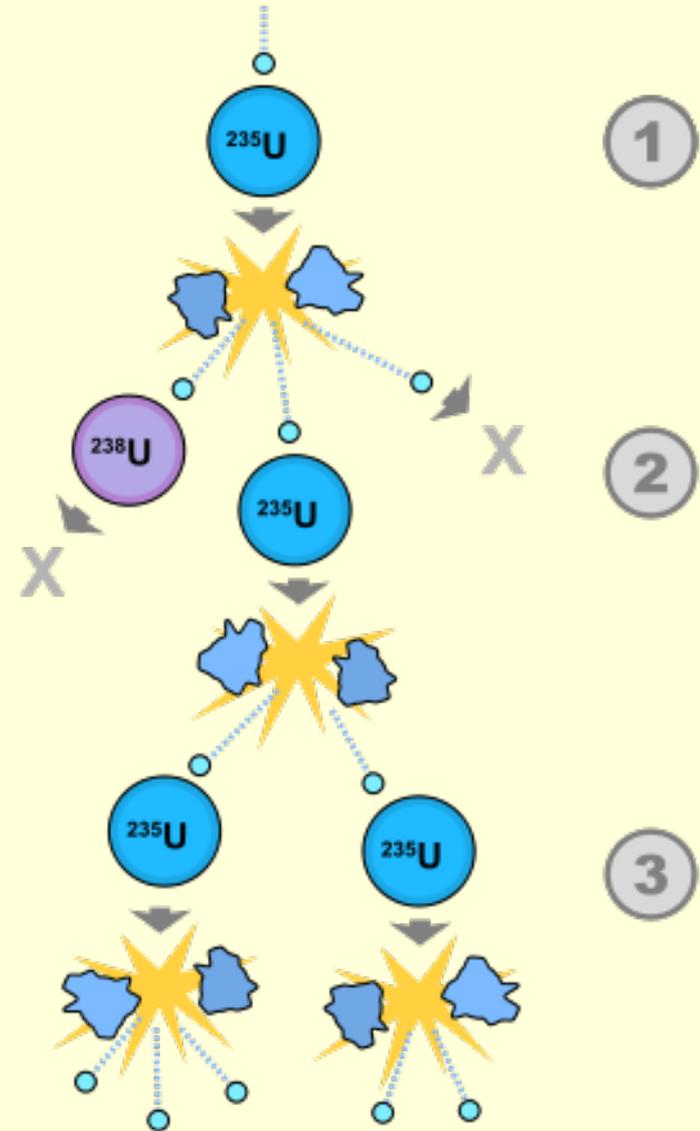
v = neutron velocity, cm/sec

Nuclear Material Definitions

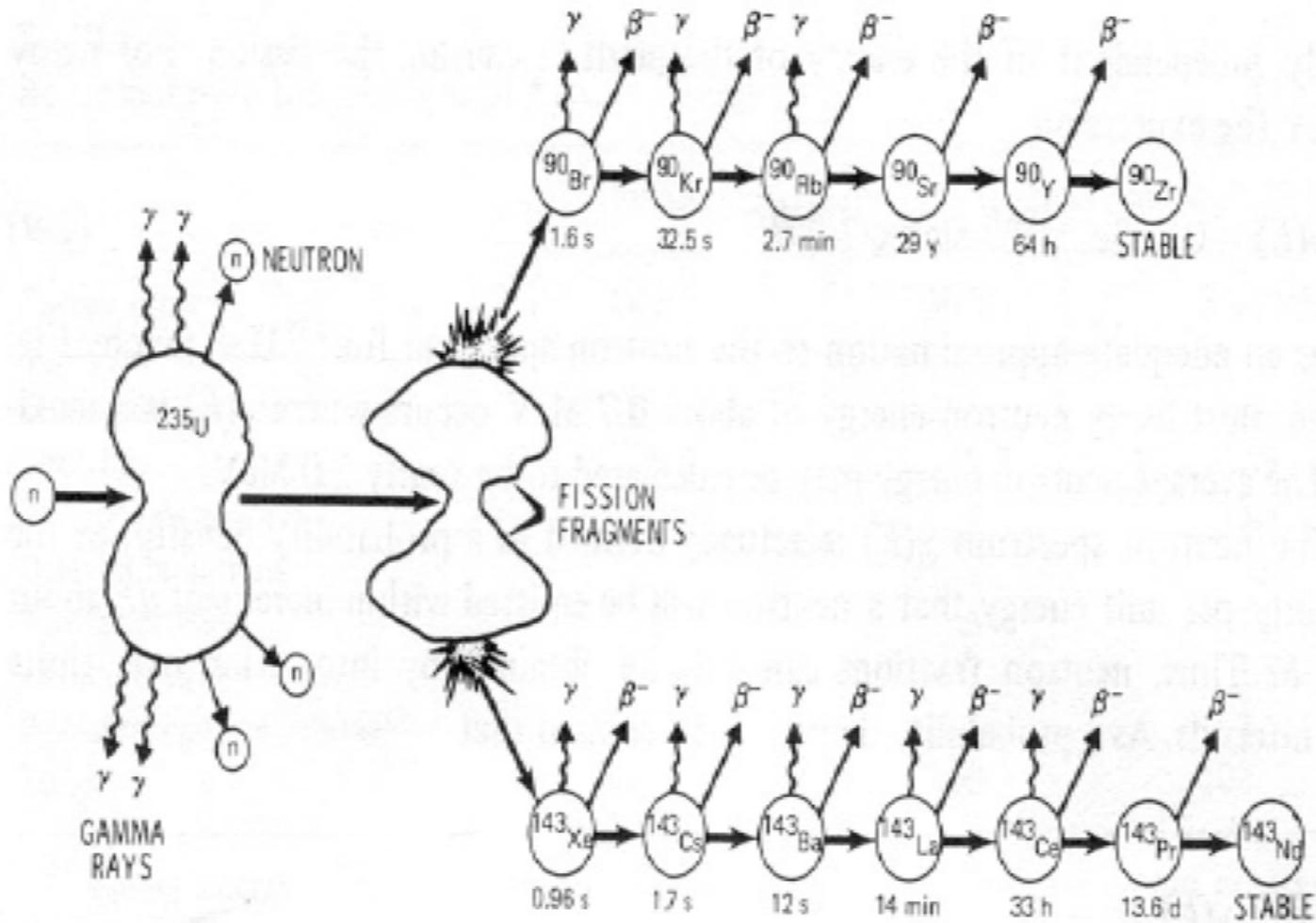
- Fissile: A nuclide that can support a self-sustaining nuclear reaction
 - ^{233}U , ^{235}U , ^{239}Pu , ^{241}Pu
 - Others which can only be made in small amounts
 - Only ^{235}U occurs naturally
- Fissionable/fissible: Nuclide can fission but not support a self-sustaining nuclear reaction
 - Virtually any actinide given high-energy neutrons
- Fertile: A nuclide that can be converted into fissile material
 - ^{238}U , ^{240}Pu , ^{232}Th

Chain Reaction

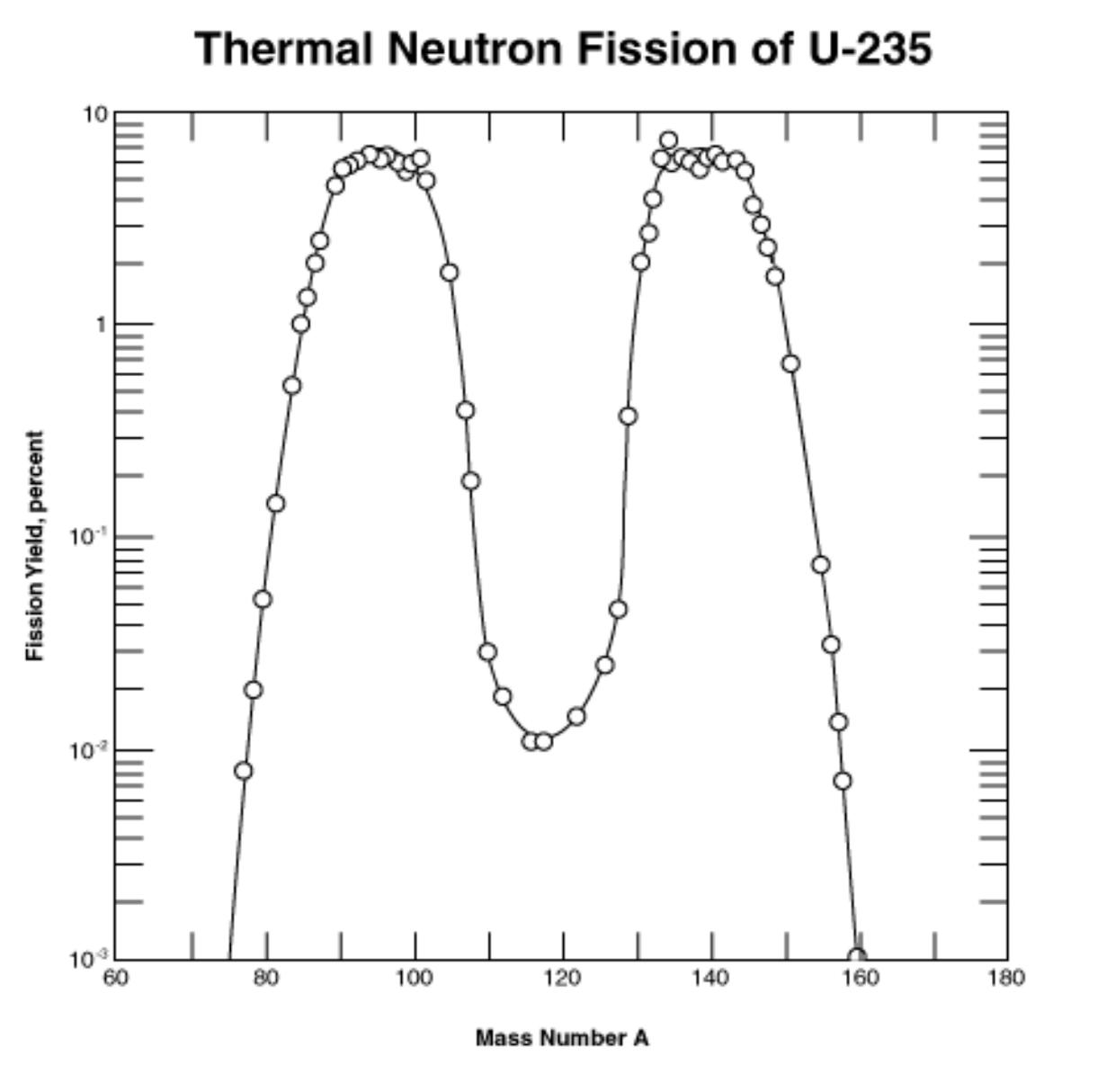
- Self-sustaining nuclear reaction



Fission Product Decay Chain-1



Fission Product Yield Curve



Binding Energy

- The mass of a nucleus is less than the sum of the masses of its protons and neutrons
- The difference is the energy that holds the nucleus together: binding energy
- Fission produces fission products and neutrons having less total binding energy
- The difference is the energy released in fission
 - Nuclear binding energy = Δmc^2

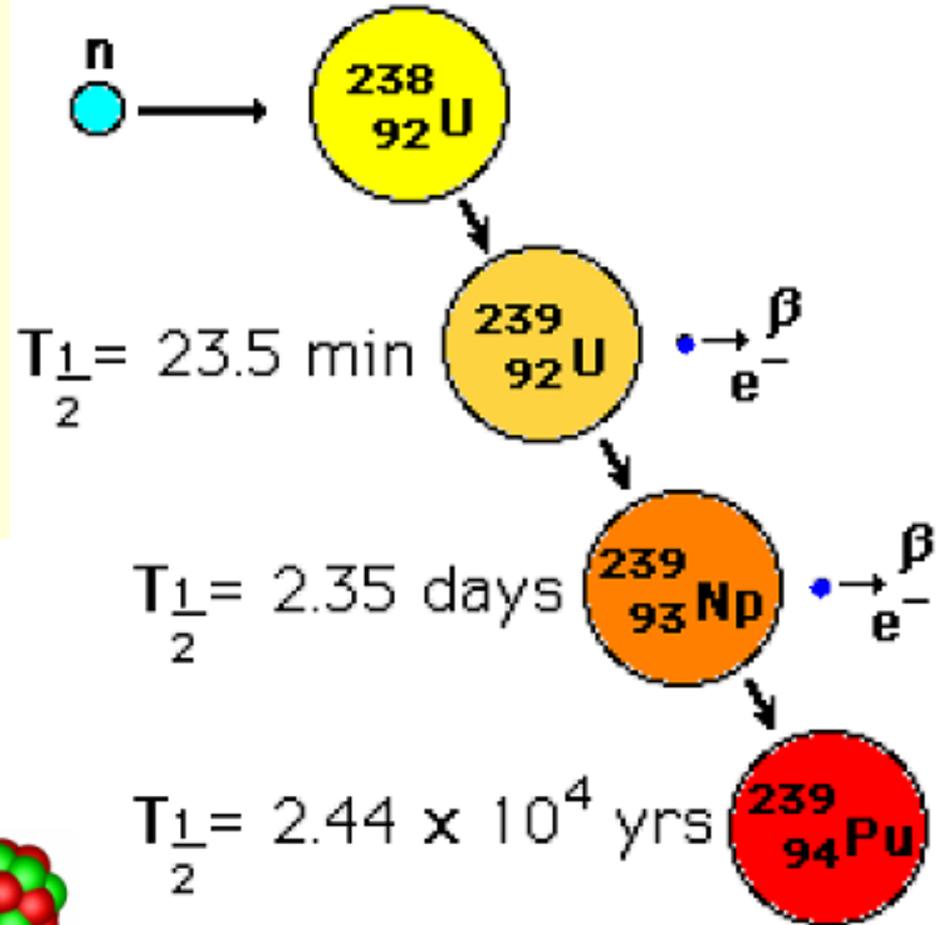
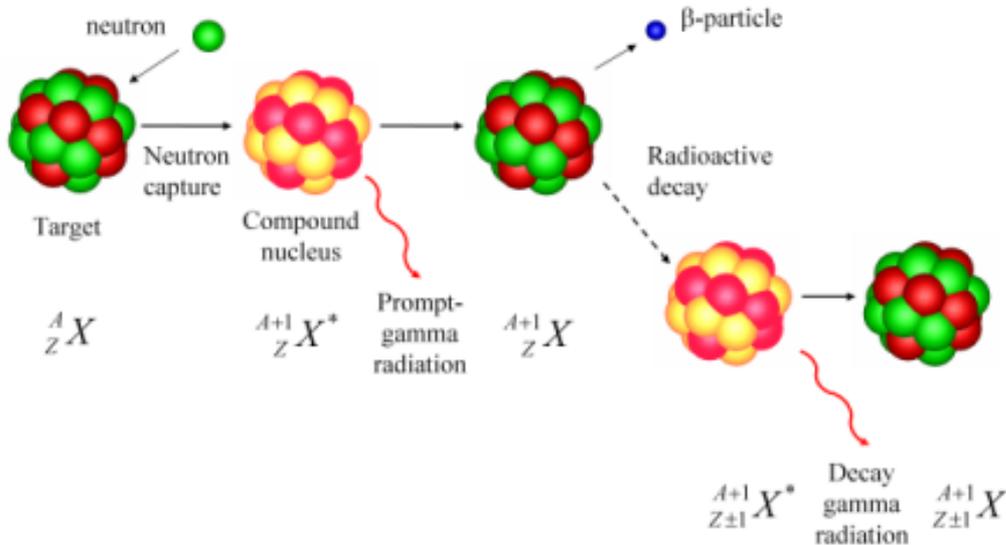
Energy per Fission

ENERGY DISTRIBUTION FOR FISSION INDUCED
BY THERMAL NEUTRONS IN ^{235}U

<i>Source</i>	<i>Energy (MeV)</i>
Fission product kinetic energy	168
Neutron kinetic energy	5
Fission γ 's (instantaneous)	5
Fission γ 's (delayed)	6
Fission product β 's	7
Total available as heat	<u>191</u>
Neutrino energy (not available as heat)	11
TOTAL	<u>202</u>

Neutron Capture: n,γ

- Addition of one neutron to a nuclide without fission
- Parasitic or productive



Neutron Capture: Other

- Neutron capture can lead to results other than gamma rays
 - Neutrons out: $(n, 2n)$; $(n, 3n)$
 - Charged particles out: (n, p) , (n, α)
 - Tend to require high-energy neutrons and have small cross sections
 - Exception: When capture product is magic
- Notation: Absorption cross section is the sum of all neutron capture and fission reactions

Neutron speed, energy, temperature

- $E(\text{J}) = 0.5 * 1.6755 \times 10^{-27} \text{ kg} * (2200 \text{ m/s})^2$

$$E(\text{J}) = 4.055 \times 10^{-21} \text{ J}$$

- $E(\text{eV}) = 4.055 \times 10^{-21} \text{ J} * 6.24 \times 10^{18} \text{ eV/J}$

$$E(\text{eV}) = 0.0253 \text{ eV}$$

Boltzman Constant

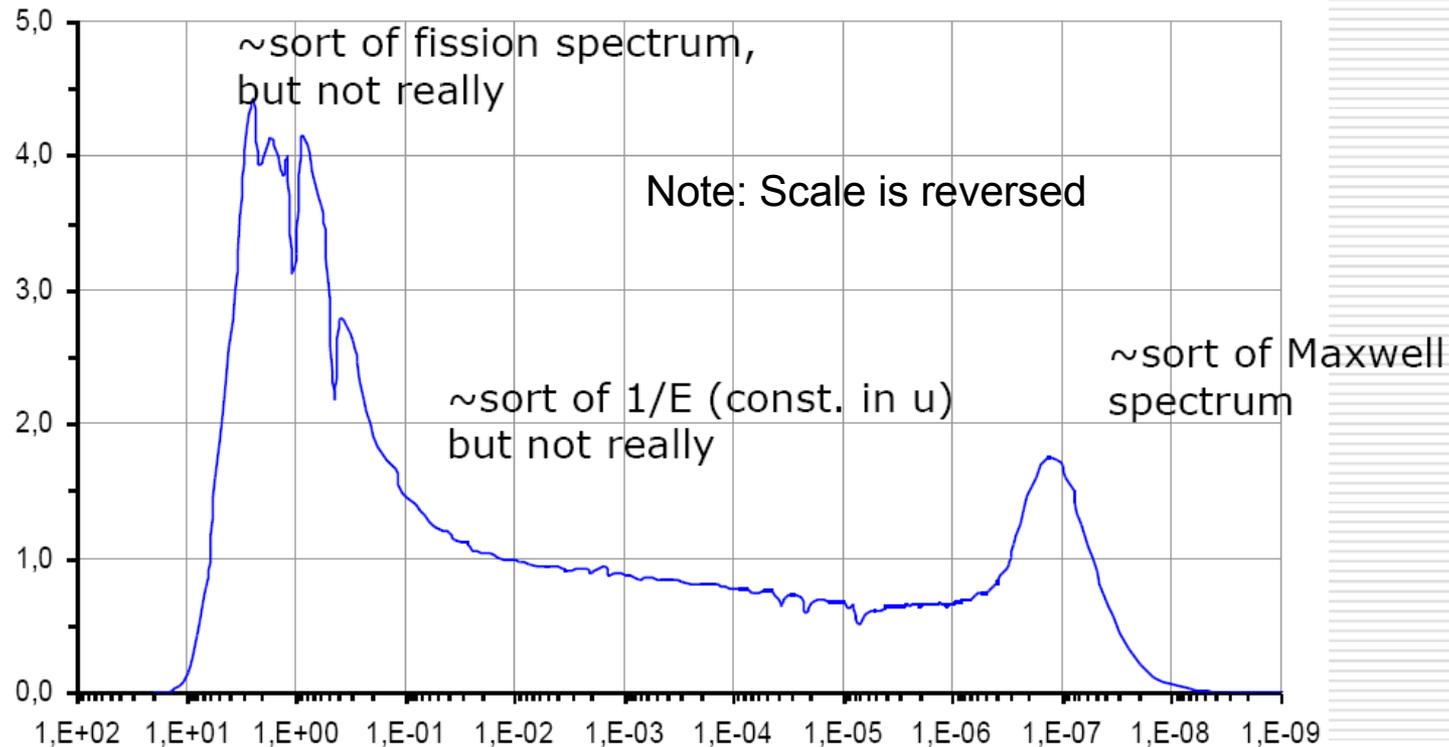
- $E(\text{T}) = 4.055 \times 10^{-21} \text{ J} / 1.38 \times 10^{-23} \text{ J/K} = 294 \text{ K}$

– $294 \text{ K} = 21 \text{ C} = 70 \text{ F}$

– Thermal neutron

Range of Neutron Spectrum

- Neutron energy spectrum can range from fraction of an eV to ~15 MeV in a reactor



$u=0$

E en MeV

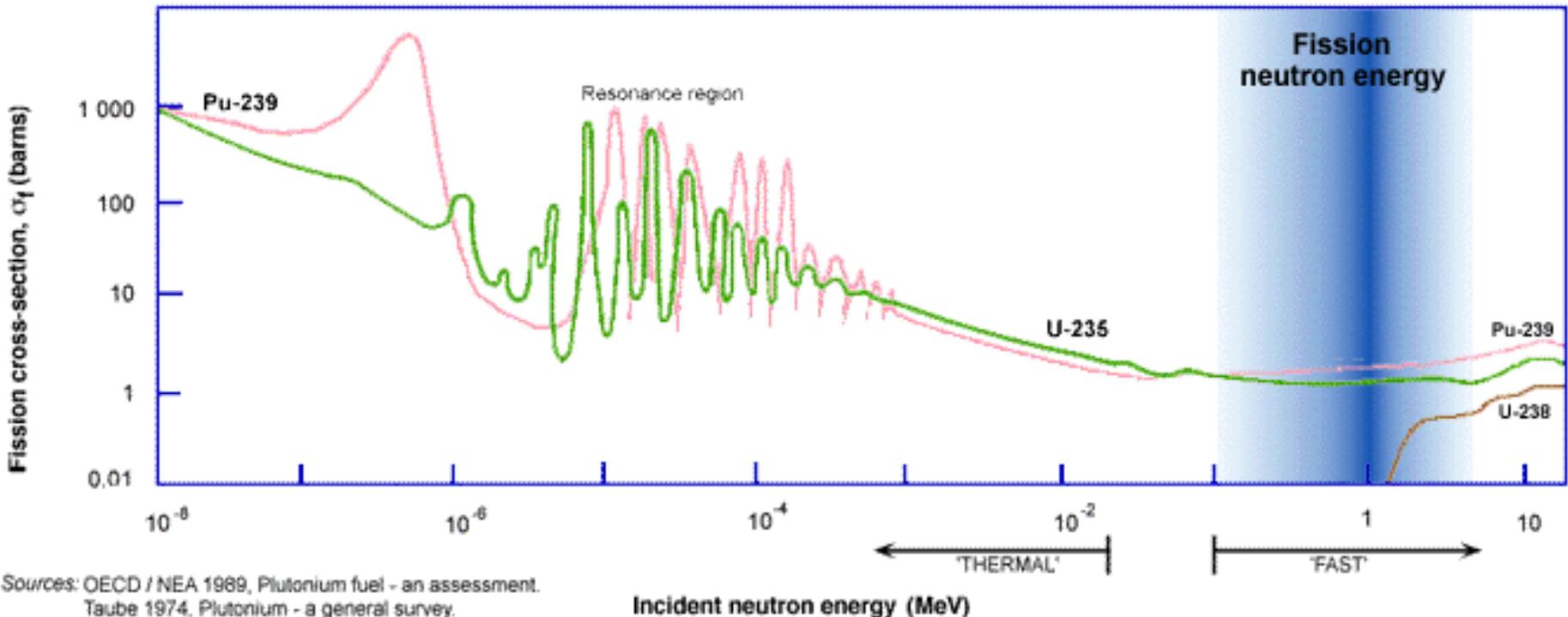
$u=11$

$=\log(1E2/1E-9)$

Typical Cross Sections

- Complication: Cross sections often vary by orders of magnitude across the neutron spectrum

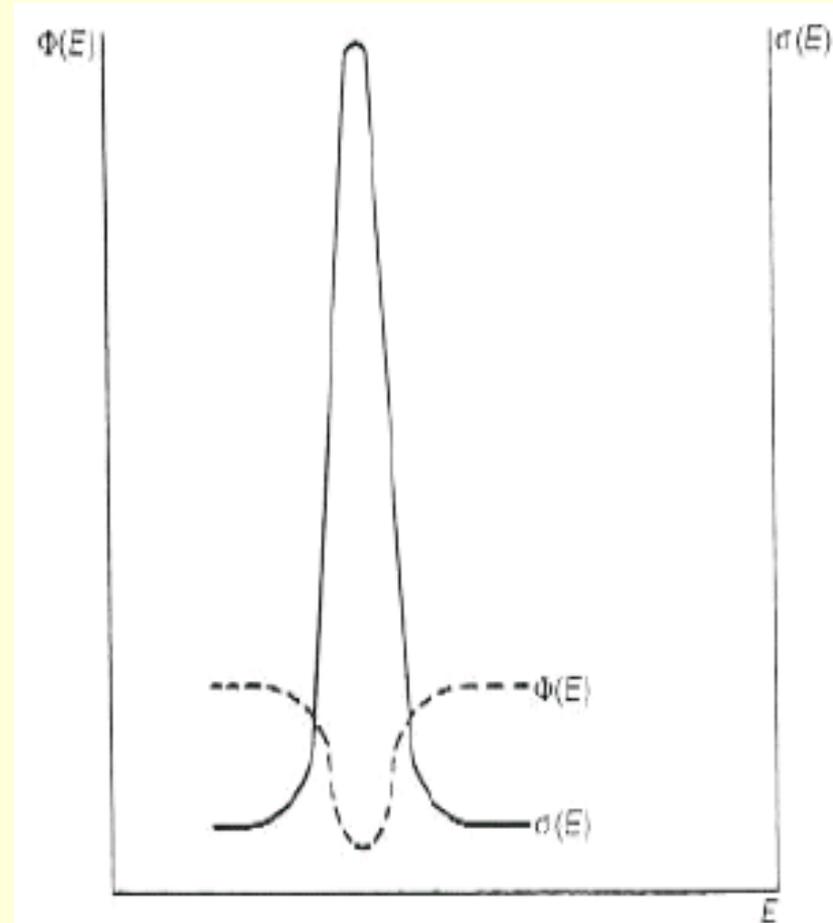
NEUTRON CROSS-SECTIONS FOR FISSION OF URANIUM AND PLUTONIUM



Sources: OECD / NEA 1989, Plutonium fuel - an assessment.
Taube 1974, Plutonium - a general survey.
1 barn = 10^{-28} m², 1 MeV = 1.6×10^{-13} J

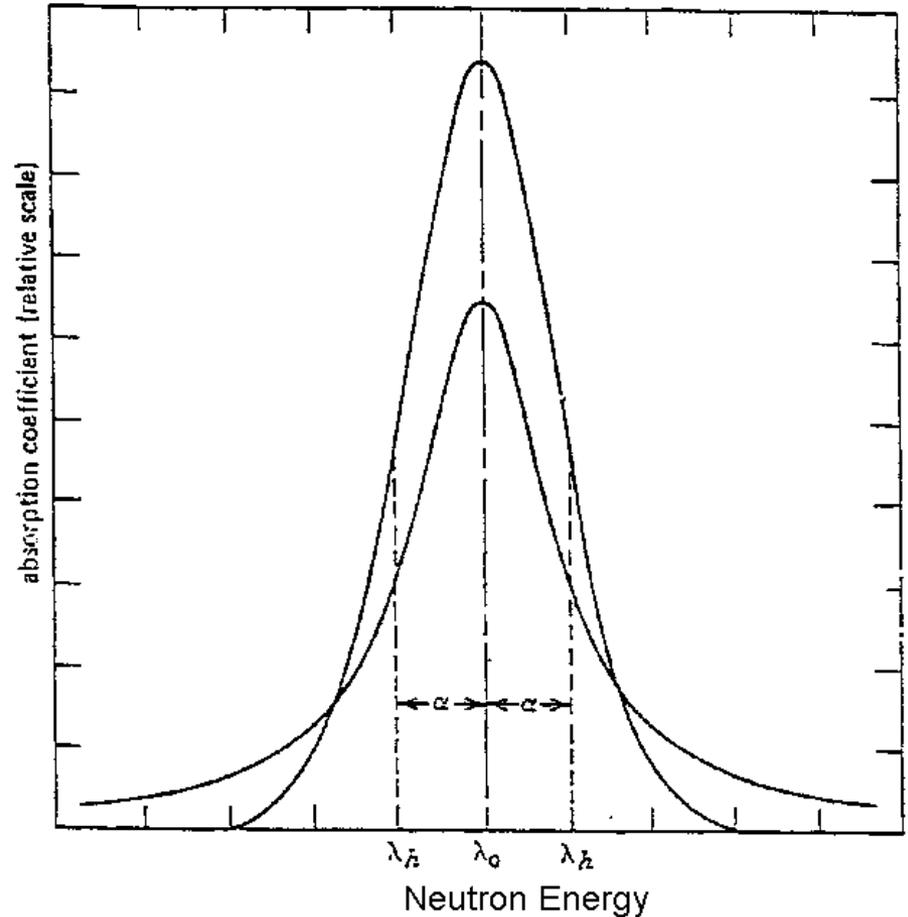
Effective Size of a Resonance

- Complication:
Resonance self shielding
- Resonances deplete supply of neutrons at a particular energy
 - In effect, resonances are smaller than measured
 - Effect can occur because resonance is large and/or concentration of nuclide is high



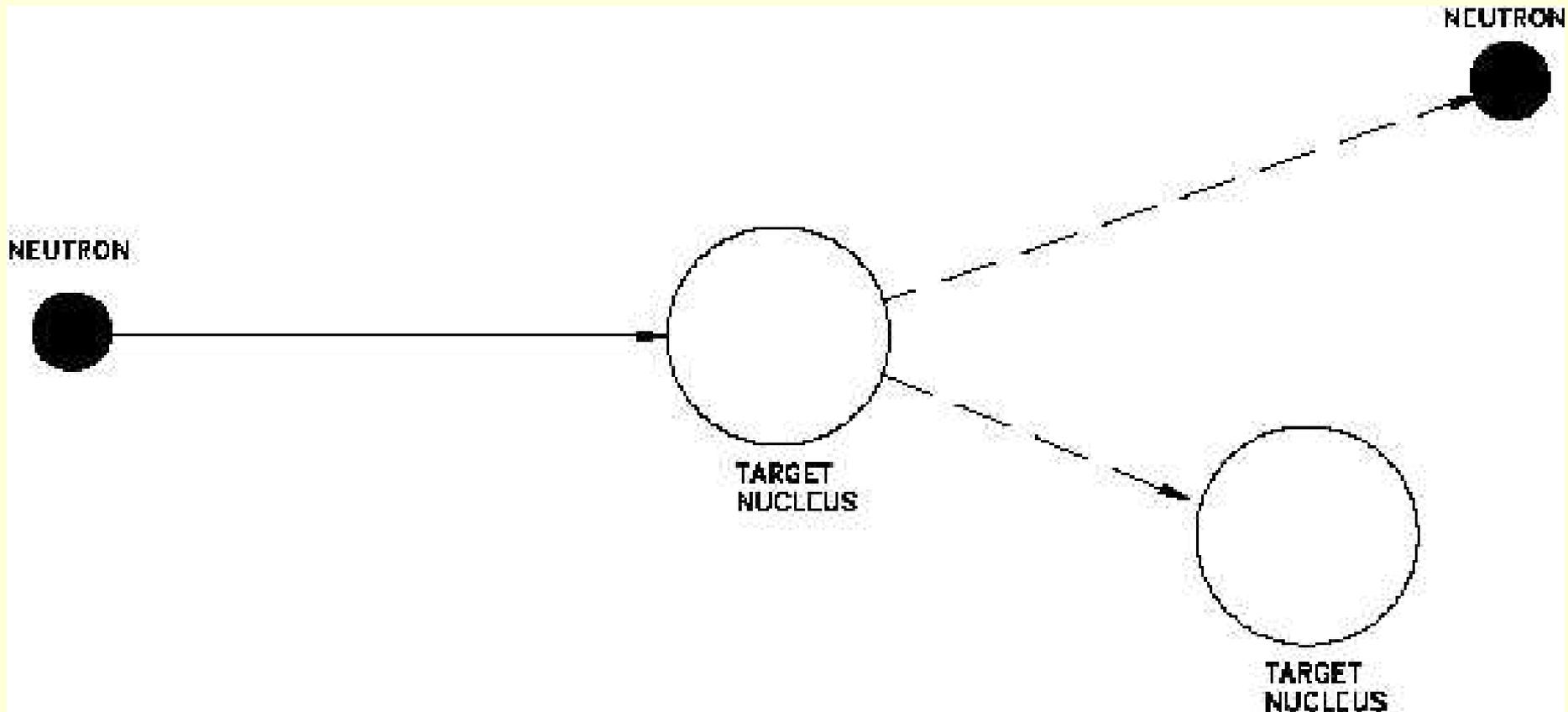
Doppler Broadening

- Resonance data measured at room temperature
- Complication: resonances get shorter and wider as temperature increases
 - Net effect is to increase size of the resonance



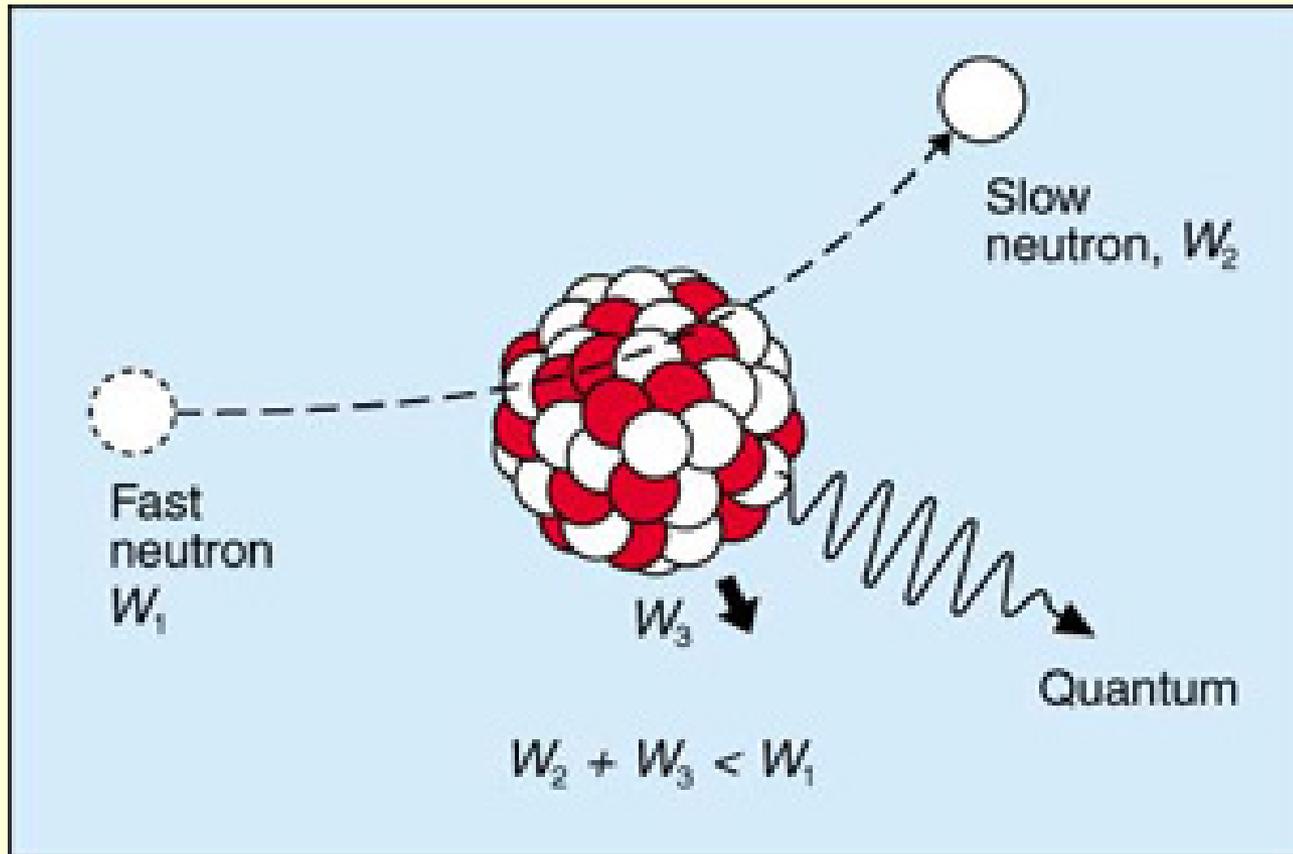
Neutron Scattering: Elastic

- Elastic neutron scattering: no change in total kinetic energy of particles
 - ‘Billiard ball’ model



Neutron Scattering: Inelastic

- Kinetic energy of final particles is less than that of initial particles
 - Energy difference released as gamma rays



Neutron Moderation

- Complication: For many reactors it is desirable for the cross section to be as large as possible
 - $R = N^* \sigma^* \varphi = \Sigma^* \varphi$
 - Reduces the required concentration of nuclear material in the fuel
- Need to slow neutrons to thermal energies
 - Notation: Neutron moderation
- Moderation occurs by neutron scattering off moderator nuclei

Neutron Moderation

- Criteria for a good neutron moderator
 - Maximizes neutron scattering
 - Minimizes non-productive neutron capture
 - Has a low atomic mass
 - Maximum neutron energy loss per collision is proportional to

$$1 - \left(\frac{A - 1}{A + 1} \right)^2$$

which is at a maximum when $A=1$ and drops to 0.28 for $A=12$

Neutron Moderation

- Moderating ratio takes criteria into account

Moderator	Moderating Ratio
Water	58
Heavy water	21,000
Helium, 1 atm	10^{-5}
Beryllium	130
Graphite	200

Thermal Reactor Criticality and Control

What is a thermal reactor

- A nuclear reactor in which the neutrons are moderated and most fissions are caused by thermal neutrons
 - Thermal neutrons = neutrons in thermal equilibrium (~Maxwellian distribution) at the temperature of their environment

Thermal Neutron Cycle

Step in Neutron Cycle	Factor	Change in number of neutrons	Total number of neutrons	Notation
Start: 1000 fission neutrons			1000	
Fast fission factor- mainly U-238	1.044	44.0	1044 ϵ	
Fast neutron non-leakage probability	0.861	-145.1	899 P_1	
Epithermal neutron non-leakage probability	0.952	-43.1	856 P_2	
Epithermal neutrons escaping U-238 capture	0.816	-157.5	698 p_{28}	
Epithermal neutrons escaping non-fuel capture	0.970	-20.9	677 p_c	
Epithermal fissions in U-235	1.034	23.7	701 B	
Thermal neutron non-leakage probability	0.985	-10.5	691 P_3	
Thermal neutrons absorbed in fuel (utilization)	0.817	-126.4	564 f	
Reduced absorp because of fuel temp	0.990	-5.6	559 T	
Net fission neutrs per neut absorbed	1.790	441.3	1000 η	

Getting to Critical

- First order effects
 - Increase concentration of fissile material
 - Decrease parasitic neutron absorption
 - Select materials with low cross sections
 - Less material in the reactor
 - Reduce neutron leakage
 - Neutrons leak from reactor surface
 - Larger reactors have lower surface-to-volume ratio

Reactor Power

- Power (watts) = $8.3 \times 10^{10} \cdot \sigma_f \cdot M \cdot \varphi$

where

- σ_f = Fission cross section, cm^2
- M = Mass of fissile material, g
- φ = Neutron flux, neutrons/ cm^2 -sec

Reactor Control

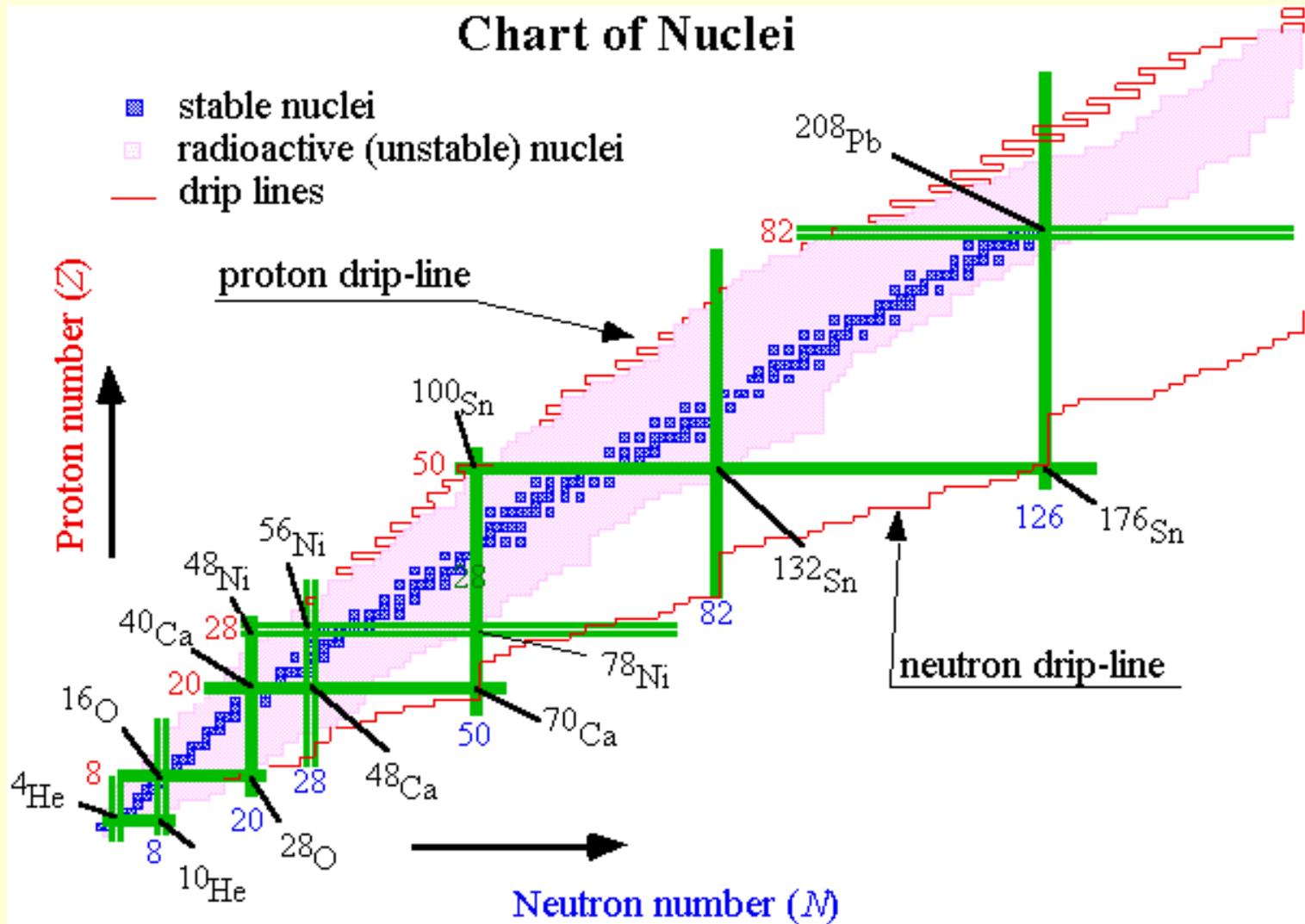
Prompt Neutrons

- On average, each fission releases 2.5 to 3 fast neutrons within 10^{-13} seconds
- The time from one generation of prompt neutrons to the next is $\sim 10^{-5}$ seconds
- Stable period: the time it takes to increase the neutron flux (and reactor power) by e (2.72x)
- For prompt neutrons the stable period is a fraction of a second
- Under these conditions a reactor would be unstable and uncontrollable

Delayed Neutrons

- A small fraction of fission products decay by emitting neutrons
- These neutrons are emitted at lower energy than prompt neutrons and with a defined half-life along the neutron drip line
- The existence of decay neutrons increases the stable period to several seconds which allows the reactor to be controlled

Neutron Drip Line



Nuclear Reactor Control

- Neutron poison: A non-productive neutron absorbing material typically having a large neutron capture cross section
 - More accurately “fission poison”
- Approach: Vary the amount of neutron absorber (i.e., the amount of non-productive neutron absorption) to achieve stable operation or slowly changing power levels

Neutron Poisons

Material	Macroscopic Thermal Absorption Cross Section, barns
Boron	107
Silver	4
Cadmium	113
Indium	7
Samarium	155
Europium	90
Gadolinium	1400

Inherent Control Mechanisms

- Increased temperature normally decreases neutron interaction
 - Doppler broadening: more neutrons absorbed in resonances
 - Thermal expansion of fuel and core: more neutron leakage
 - Boiling of coolant: less moderation
 - This effect can increase neutron interactions in fast reactors under some circumstances

Reactor Physics Calculations

Definition

- The science of the interaction of elementary particles and radiations characteristic of nuclear reactors with matter in bulk

What Needs to be Known?

- Neutron flux in and around a reactor core, and interactions and reactions of neutrons and other radiation in fuel and structural materials as a function of
 - Space
 - Energy
 - Direction
 - Time

How It is Done

- Solve the Boltzmann equation describing transport of neutral particles
 - Essentially the same approach for neutrons and photons

Boltzman Equation

$$\begin{aligned} \frac{1}{v(E)} \frac{d\Phi(\mathbf{r}, E, \boldsymbol{\Omega}, t)}{dt} & \textcircled{1} \\ & = -\boldsymbol{\Omega} \cdot \nabla \Phi(\mathbf{r}, E, \boldsymbol{\Omega}, t) \textcircled{2} - \Sigma_t(\mathbf{r}, E, \boldsymbol{\Omega}) \Phi(\mathbf{r}, E, \boldsymbol{\Omega}, t) \textcircled{3} \\ & \quad + \chi(E) \int_{E'} dE' \int_{\Omega'} d\Omega' \nu \Sigma_f(\mathbf{r}, E', \boldsymbol{\Omega}') \Phi(\mathbf{r}, E', \boldsymbol{\Omega}', t) \textcircled{4} \\ & \quad + \int_{E'} dE' \int_{\Omega'} d\Omega' \Sigma_s(\mathbf{r}; E' \rightarrow E; \boldsymbol{\Omega}' \rightarrow \boldsymbol{\Omega}) \Phi(\mathbf{r}, E', \boldsymbol{\Omega}', t) \end{aligned}$$

Term 1 = rate of accumulation of neutrons

Term 2 = rate of leakage out of the element

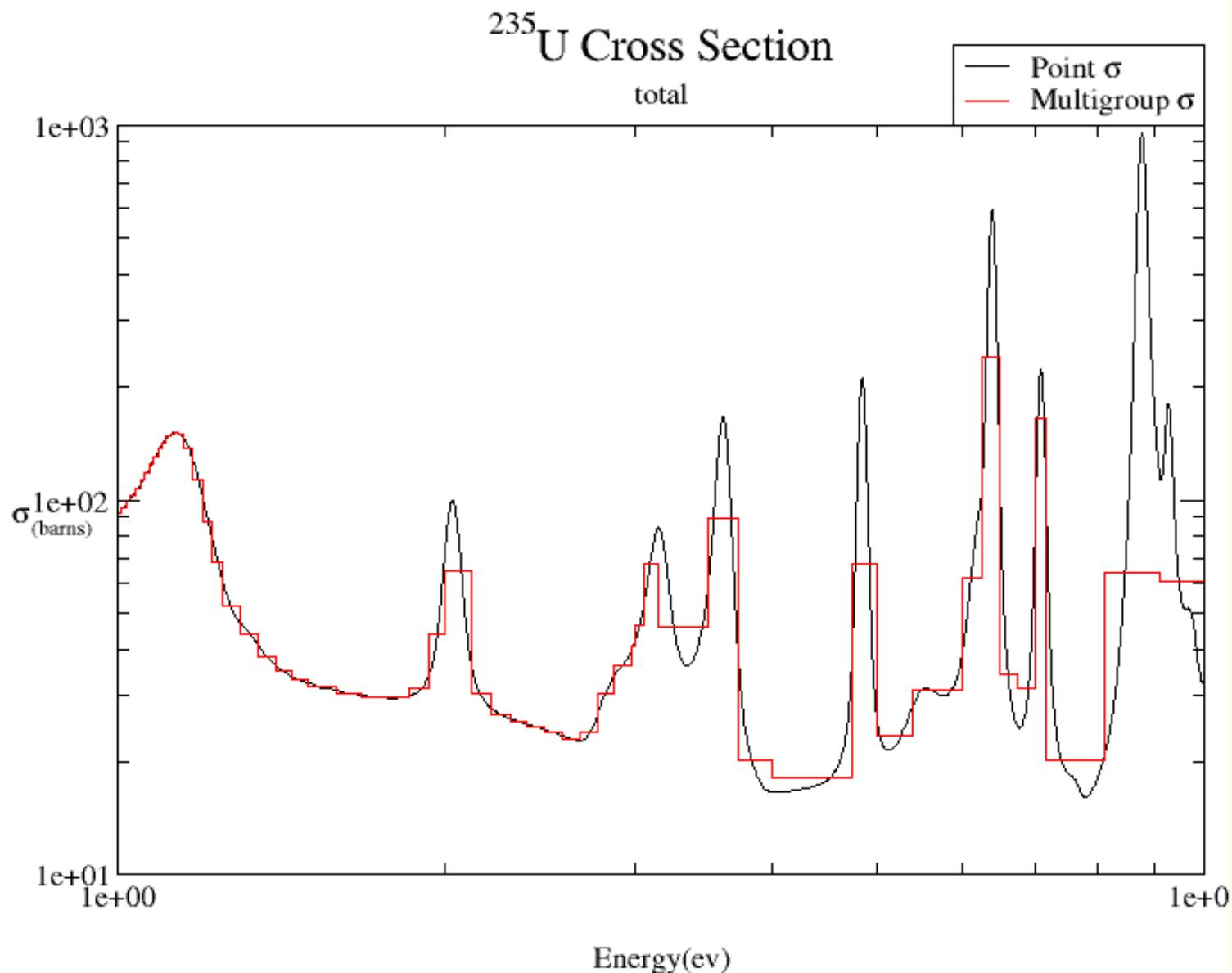
Term 3 = total interaction rate (removal due to absorption or scatters out of volume element or out of energy, E).

Terms 4 and 5 represent production phenomena where neutrons at E' and $\boldsymbol{\Omega}'$ react with nuclei to generate neutrons at E and $\boldsymbol{\Omega}$. Integrals sum over all initial energies and direction.

Term 4 = total fission rate; then $\chi(E)$ refers to the energy spectrum produced in the fission process.

Term 5 = differential scattering of neutrons from initial energy, E', to final energy, E, and from initial direction $\boldsymbol{\Omega}'$, to final direction, $\boldsymbol{\Omega}$. The cross section, $\Sigma_s(\mathbf{r}; E' \rightarrow E; \boldsymbol{\Omega}' \rightarrow \boldsymbol{\Omega})$ accounts for the relative probabilities of all possible combinations.

Step 1: Multigroup Approximation

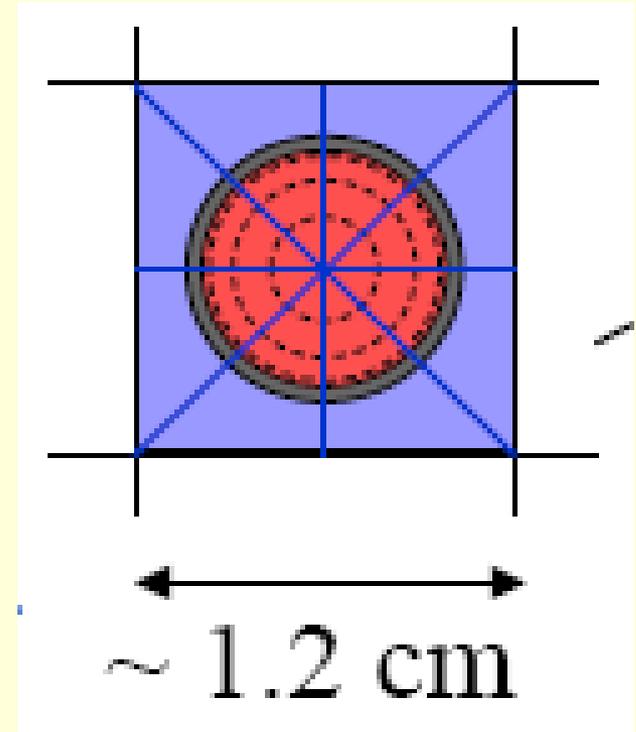


Boltzman Solution Complication

- Accounting for heterogeneity of reactors: with variation in space, energy, neutron angle, time the number of points is in the billions

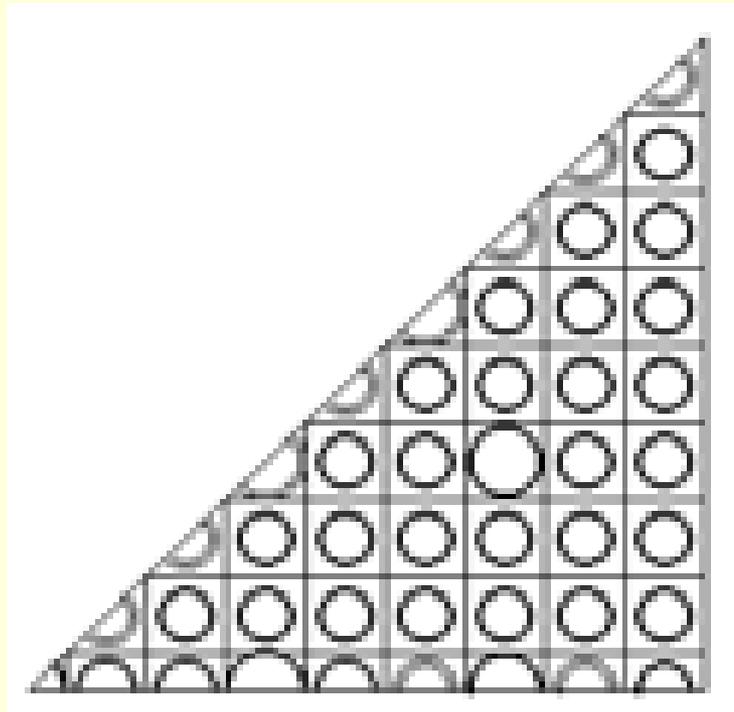
Step 1: Start Small with Detail

- Start with a single fuel rod
- Using transport or Monte Carlo methods and multigroup cross sections calculate the many group neutron flux
 - Static
 - Two dimensions
 - Only key nuclides, lumped fission products
- Done for a number of fuel compositions



Step 2: Start Small with Detail

- With modern computing power it may be possible to do many-group modeling of a number of symmetrical cells in an array

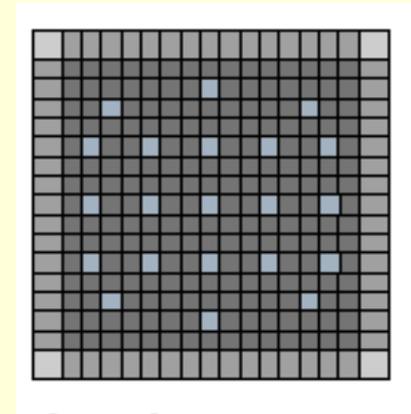


Step 3: Homogenize and Grow

- Use the many group flux to weight the many group cross sections to yield few (2 to 5) group cross sections
- Reaction rates (product of flux and cross section are preserved) during homogenization



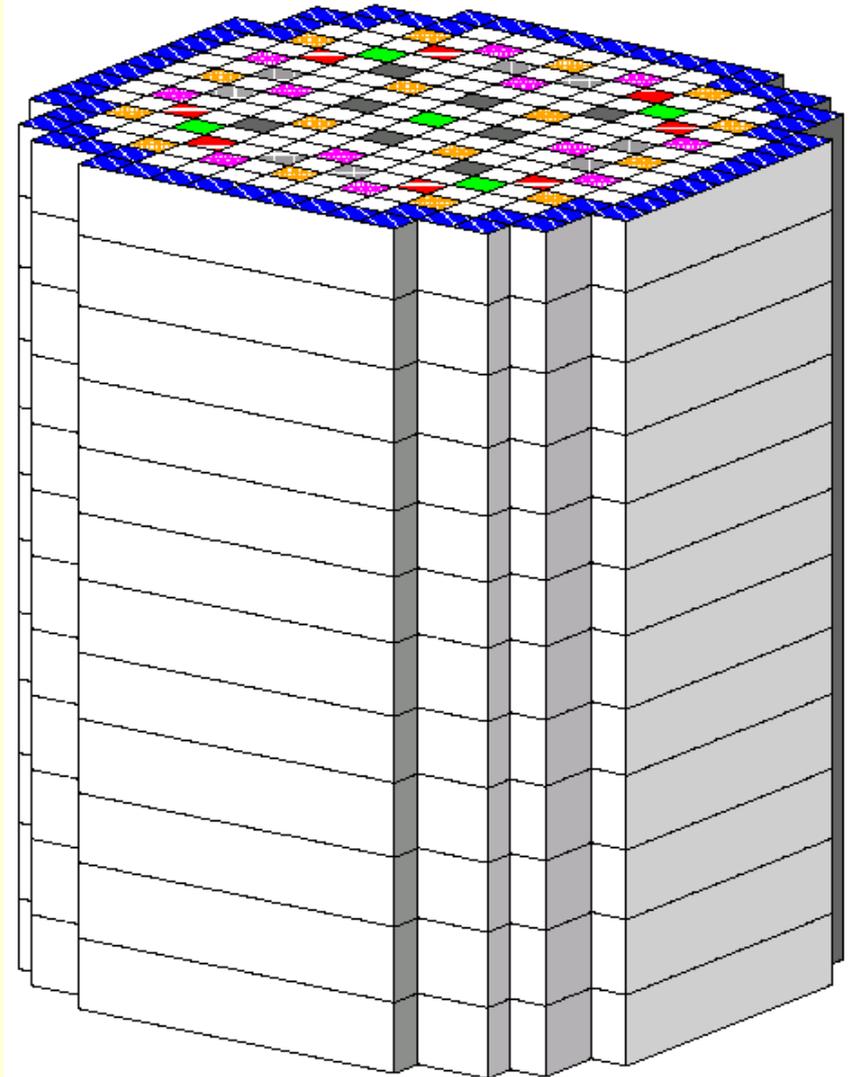
Array



Cells

Step 4: The Whole Core

- 3-D model of entire core
 - Time dependent
 - Few energy groups
 - Relatively coarse grid
- Stepwise iteration with depletion (next slide)
- Diffusion theory is fast and sufficiently accurate when homogenized
- 3-D codes are run many times to optimize fuel composition and movement



Step 5: Detailed Depletion

- Use few group fluxes to weight cross sections for many nuclides (hundreds) to yield total flux and spectrum-averaged (1-group) cross sections
- Use cross sections and flux in models to predict the buildup and decay of many radionuclides as a function of time
 - Simple conversions yield other properties such as gamma ray intensity and decay heat
- Boltzman equation not needed: solve the Bateman equations

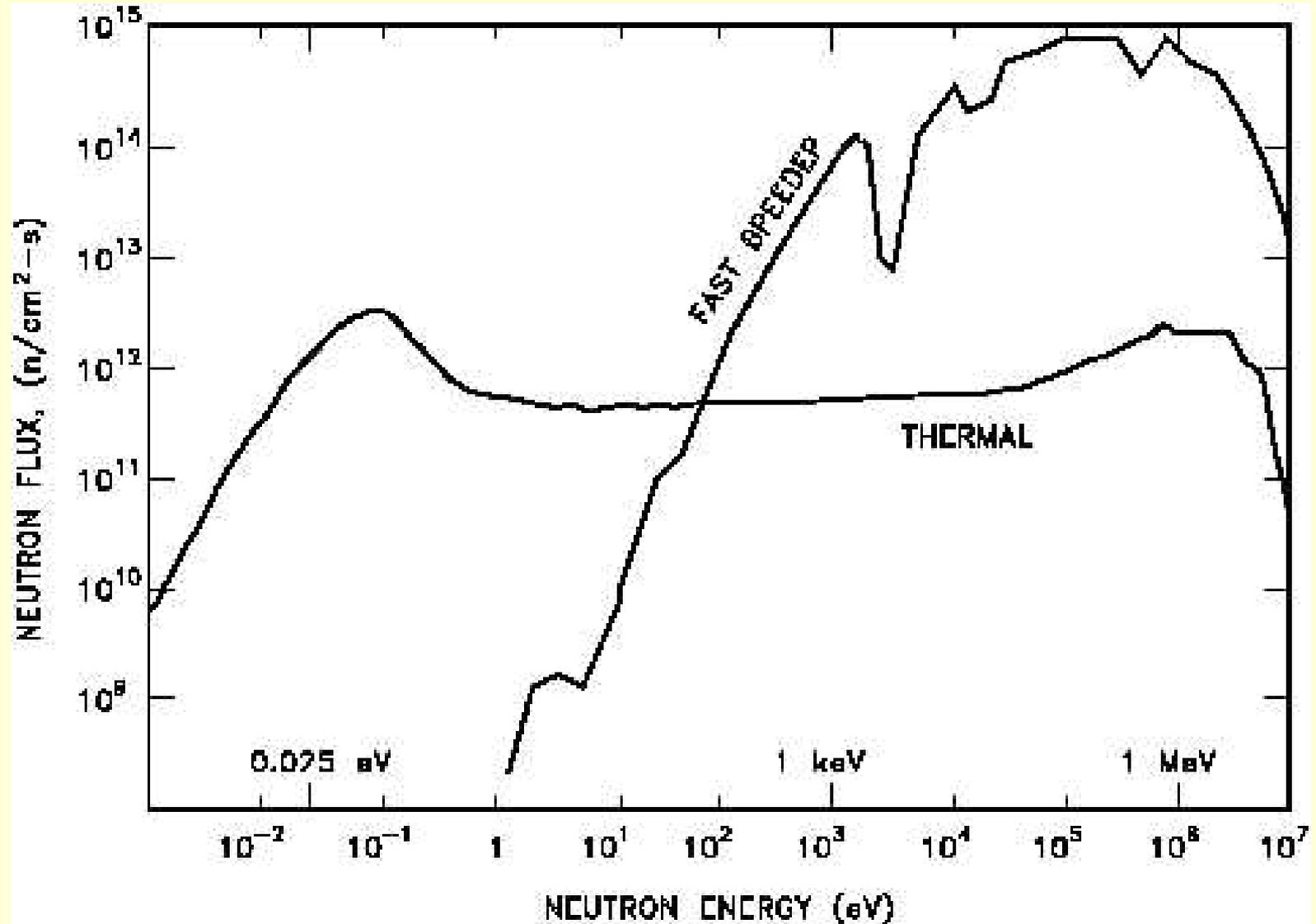
Bateman Equation

- The Bateman equation describes the buildup and decay of nuclides given knowledge of flux and cross sections
 - It can be solved in closed form for many cases
 - Closed form is tedious but computer codes alleviate this
 - Complication: Closed form not possible when a radionuclide produces itself
 - Common for the actinides because of alpha decay
 - Solution: Solve equation numerically
 - Computer codes exist; results later

$$\frac{dN_i}{dt} = -\sum_{j \neq i} \left[\lambda_{ji}^d + \int \varphi(E, t) \sigma_{ji}^{tr}(E) dE \right] N_i + \sum_{j \neq i} \left[\lambda_{ij}^d + \int \varphi(E, t) \sigma_{ij}^{tr}(E) dE \right] N_j$$

Fast Reactor Physics

Fast Reactor Neutron Spectrum



Ramifications of Fast Spectrum-1

- Criticality considerations simpler
 - No need to consider thermal region
 - Resonances are much less important
- But
 - Fissions in fissionable nuclides (e.g, U^{238}) are more important

Ramifications of Fast Spectrum-2

- Cross sections are smaller at higher energy because of general $1/v$ dependence
 - Need higher concentration of fissile material to achieve criticality
 - Neutron fluxes are higher to achieve desired power level

Ramifications of Fast Spectrum-3

- Higher ratio of fission-to-absorption cross sections
 - Fewer neutron losses to unproductive capture
 - Convert fertile nuclides to fissile nuclides faster than they consume fissile nuclides: they can be “breeders”
 - Less production of minor actinides: Np, Am, Cm

